

(A1) For each particle: $\underline{F}_i = \dot{\underline{p}}_i$ where $\underline{p}_i = m_i \dot{\underline{r}}_i$ [1]

Summing: $\sum_i \underline{F}_i = \sum_i \dot{\underline{p}}_i = \dot{\underline{P}}$ where $\underline{P} = \sum_i \underline{p}_i$ [2] define \underline{P} [2]

$\sum_i \underline{F}_i = \sum_i \underline{F}_i^{\text{ext}} + \sum_i \sum_{j \neq i} \underline{F}_{ij}$

$\therefore \sum_i \underline{F}_i = \sum_i \underline{F}_i^{\text{ext}} = \underline{F}^{\text{ext}}$ falls apart into sum over all pairs $\underline{F}_{ij} + \underline{F}_{ji} = 0$. [1]

and $\underline{F}^{\text{ext}} = \frac{d\underline{P}}{dt}$ (*) [1] [4]

(A2) IF the rotational motion takes place about an axis that traces the surface of a cone when it bodily turns, then the axis is said to precess around the vertical direction and the phenomenon is called precession. [1] [3]

(Example: motion of a spinning top.)

(A3) Just relate two expressions for grav. force: [2]

$mg = \frac{GMm}{r_e^2} \rightarrow \underline{g} = \frac{GM}{r_e^2}$ [1] [3]

(A4) Force is central [1] $\underline{F} = f(r) \underline{e}$ $\therefore \frac{d\underline{L}}{dt} = \underline{e} \times \dot{\underline{p}} = 0 \Rightarrow \underline{L} = \text{const.}$ [2]

But \underline{L} always perp. to plane defined by \underline{e} and \underline{v} , hence whole motion lies in a plane. [2] [5]

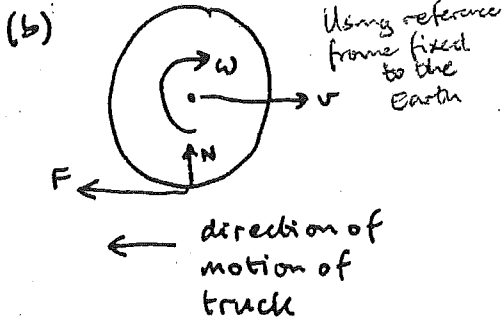
(A5) [2] $-\underline{\omega} \times (\underline{\omega} \times \underline{R})$ $\underline{\omega}$ = Earth's angular velocity [1]
 \underline{R} = Earth's radius (outward from center)

$\frac{\omega^2 R}{g} \approx 0.35\%$ [2] [5]

(B1) (a) let ρ be mass per unit cross-sectional area:

Then, [2] $m = \int_0^a \rho 2\pi r dr = \pi \rho a^2 \rightarrow$ sum over thin cylindrical shells.

[2] $I = \int_0^a \rho 2\pi r \cdot r^2 dr = \frac{\pi \rho a^4}{2} = \frac{1}{2} m a^2$ [6]



linear motion of CM: $F = -m\dot{v}$ ① [2]

ang. motion about CM: $Fa = I\dot{\omega}$ ② [2]

no-slip condition: $v - a\omega = -\frac{1}{2}gt$ ③ [2]

From ① and ②

$I\dot{\omega} = -m a \dot{v}$

$a\dot{\omega} = -2\dot{v}$

\Rightarrow

$a\omega = -2v$ [2]

since $\omega = 0$ when $v = 0$

use boxed result in ③ to find:

$v + 2v = -\frac{1}{2}gt \Rightarrow v = -\frac{1}{6}gt$ ④ [2]

so,

$a\omega = \frac{1}{3}gt$

$a\theta = \frac{1}{6}gt^2$ ⑤ since $\theta = 0$ at $t = 0$ [2]

↑ angle turned through

When leaves edge of truck: $a\theta = L = 5\text{m}$

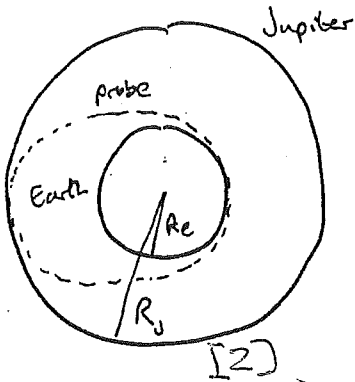
From ⑤ $t^2 = \frac{6L}{g}$

subst. in ④ $v = -\frac{1}{6}g \sqrt{\frac{6L}{g}} = -\sqrt{\frac{Lg}{6}}$

$= -\sqrt{\frac{5 \times 9.8}{6}} \text{ms}^{-1}$ [2]

or speed 2.86ms^{-1} in direction of motion of truck

(B2)



Earth's orbit, radius R_e
 Jupiter's orbit, radius $R_j = 5.2 R_e$

(i) For earth's orbit: [2]

$$\frac{v_e^2}{R_e} = \frac{GM}{R_e^2} \quad [16]$$

(ii) For the elliptical orbit of the probe.

Ang. mom: $L = m r^2 \dot{\theta} \Rightarrow \dot{\theta} = \frac{L}{m r^2} \dots \textcircled{1} \quad [2]$

Energy: $E = \frac{1}{2} m \dot{r}^2 + \frac{1}{2} m r^2 \dot{\theta}^2 - \frac{GMm}{r^2} \dots \textcircled{2} \quad [2]$

Equate energy at $r_{min} = R_e$ and $r_{max} = R_j$ using $\dot{r} = 0$ at those points:

$$\frac{1}{2} \frac{L^2}{m R_e^2} - \frac{GMm}{R_e} = \frac{1}{2} \frac{L^2}{m R_j^2} - \frac{GMm}{R_j} \quad [2]$$

$$\frac{L^2}{2m} \left(\frac{1}{R_e^2} - \frac{1}{R_j^2} \right) = GMm \left(\frac{1}{R_e} - \frac{1}{R_j} \right)$$

$$\frac{L^2}{2m} = \frac{GMm R_e R_j}{R_e + R_j} \quad [2]$$

At R_e , probe's speed is v and $L = m v R_e$

$$\Rightarrow v^2 = \frac{L^2}{m^2 R_e^2} = 2 \frac{GM}{R_e} \frac{R_j}{R_e + R_j}$$

$$= 2 v_e^2 \frac{R_j}{R_e + R_j} \quad [2]$$

$1.1 \sqrt{\frac{2}{1+1/5.2}} > \sqrt{2}$
 so will escape

$$\therefore \frac{v}{v_e} = \sqrt{\frac{2 R_j}{R_e + R_j}} = \sqrt{\frac{2 \times 5.2}{6.2}} = 1.3 \quad [2]$$

or note $\frac{v_0}{v_e} = \sqrt{\frac{2}{1 + \frac{R_e}{R_j}}}$ so $\frac{v}{v_e} = \sqrt{2}$ if $R_j = \infty$

If initial speed is $1.1 v$, then energy is,

$$E = \frac{1}{2} m (1.1 v)^2 - \frac{GMm}{R_e} = \frac{GMm}{R_e} \left[(1.1)^2 \frac{R_j}{R_e + R_j} - 1 \right] \quad [4]$$

$$= \frac{GMm}{R_e} (0.01) > 0 \quad \left[\begin{array}{l} \text{so probe will} \\ \text{escape} \end{array} \right] \quad [2]$$

B3 (a) Starting from: $\ddot{x} = g^* - 2\omega \times \dot{x}$

Backwards

with $x = a$ and $\dot{x} = v$ at $t = 0$

Integrate once, $\dot{x} = g^* t - 2\omega \times x + \text{const.}$ [2]

use initial conditions: $\dot{x} = v + g^* t - 2\omega \times (x - a)$ [2]

To order ω can substitute x to order 1 in the Coriolis term:

$$x = vt + \frac{1}{2} g^* t^2 + a + O(\omega)$$

so: $\dot{x} = v + g^* t - 2\omega \times (vt + \frac{1}{2} g^* t^2) + O(\omega^2)$ [2]

Integrate: $x = vt + \frac{1}{2} g^* t^2 - 2\omega \times (v \frac{t^2}{2} + \frac{1}{6} g^* t^3) + \text{const.}$ [2]

Fix const by $x = a$ at $t = 0$: $\text{const} = a$

so $x = a + vt + \frac{1}{2} g^* t^2 - \frac{1}{3} \omega \times g^* t^3 - \omega \times vt^2$ [2]

(b) dropped from a tower. Choose upward vertical to be defined from g^* .

So, $a = h \hat{z}$, $v = 0$, $g^* = -g \hat{z}$ [2]

Let \hat{x} be East, \hat{y} North.

In latitude λ , $\omega = \hat{z} \omega \sin \lambda + \hat{y} \omega \cos \lambda$

$$\omega \times g^* = -\omega g \cos \lambda \hat{x}$$

Hence in components:

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ h \end{pmatrix} - \frac{1}{2} g t^2 \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} + \frac{1}{3} \omega g \cos \lambda t^3 \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$$

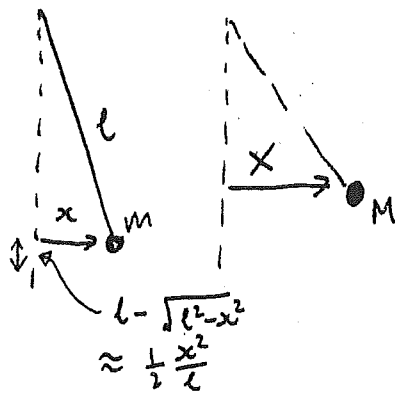
$z = 0$ when $t = \sqrt{\frac{2h}{g}}$, and then $x = \frac{1}{3} \omega g \cos \lambda \cdot \left(\frac{2h}{g}\right)^{3/2}$

$$= \frac{1}{3} \omega \left(\frac{8h^3}{g}\right)^{1/2} \cos \lambda$$

c) Eastward deflection since $x +ve$.

Normal mode: every part of system oscillates at same frequency [2]

(B4)



For small oscillations:

$$L = \frac{1}{2} m \dot{x}^2 + \frac{1}{2} M \dot{X}^2 - \frac{1}{2} g \left(\frac{x^2}{l} + \frac{X^2}{l} \right) - \frac{1}{2} k (X-x)^2$$

so E-L eqns give:

$$m \ddot{x} = -\frac{mgx}{l} + k(X-x) \quad [2]$$

$$M \ddot{X} = -\frac{MgX}{l} - k(X-x) \quad [2]$$

Look for normal mode: $x = A e^{i\omega t}$, $X = B e^{i\omega t}$

$$\begin{pmatrix} \frac{g}{l} + \frac{k}{m} & -\frac{k}{m} \\ -\frac{k}{M} & \frac{g}{l} + \frac{k}{M} \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix} = \omega^2 \begin{pmatrix} A \\ B \end{pmatrix} \quad [2] \quad [8]$$

Allowed frequencies: $\left(\frac{g}{l} - \omega^2 + \frac{k}{m}\right) \left(\frac{g}{l} - \omega^2 + \frac{k}{M}\right) - \frac{k^2}{mM} = 0$

$$\left(\frac{g}{l} - \omega^2\right) \left(\frac{g}{l} - \omega^2 - \frac{k}{m} - \frac{k}{M}\right) = 0$$

$$\omega^2 = \frac{g}{l} \equiv \omega_0^2 \quad \text{or} \quad \omega^2 = \frac{g}{l} + \frac{k}{m} + \frac{k}{M} \equiv \omega_1^2 \quad [2]$$

(i) when $\omega^2 = \frac{g}{l}$ find: $\begin{pmatrix} \frac{k}{m} & -\frac{k}{m} \\ -\frac{k}{M} & \frac{k}{M} \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix} = 0$ mode $\begin{pmatrix} 1 \\ 1 \end{pmatrix}$ [2]

(as expected) [2]

(ii) when $\omega^2 = \frac{g}{l} + \frac{k}{m} + \frac{k}{M}$ find: $\begin{pmatrix} -\frac{k}{M} & -\frac{k}{m} \\ -\frac{k}{M} & -\frac{k}{m} \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix} = 0$ mode: $\begin{pmatrix} 1 \\ -\frac{m}{M} \end{pmatrix}$ [4]

Motion starting from $\begin{pmatrix} x \\ X \end{pmatrix} = \begin{pmatrix} 0 \\ a \end{pmatrix}$, $\begin{pmatrix} \dot{x} \\ \dot{X} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$

write: $\begin{pmatrix} x \\ X \end{pmatrix} = \alpha \begin{pmatrix} 1 \\ 1 \end{pmatrix} \cos \omega_0 t + \beta \begin{pmatrix} 1 \\ -\frac{m}{M} \end{pmatrix} \cos \omega_1 t$ [2] [6] [2]

$$\Rightarrow \alpha + \beta = 0$$

$$\alpha - \frac{\beta m}{M} = a$$

$$\Rightarrow \alpha = -\beta = \frac{Ma}{m+M}, \text{ so solution is}$$

$$\begin{pmatrix} x \\ X \end{pmatrix} = \frac{aM}{m+M} \begin{pmatrix} \cos \omega_0 t - \cos \omega_1 t \\ \cos \omega_0 t + \frac{m}{M} \cos \omega_1 t \end{pmatrix}$$